



Challenges in the Understanding of Auroral Acceleration Physics

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What We Do Know – A Lot!

- ‘Mono-energetic’ electron precipitation causes the aurora.
- Parallel current plus converging field lines requires a potential drop.
- Both upward and downward currents play a role, but the two regions are different.
- Background parameters: both ionospheric and plasma sheet source populations.
- Alfvén waves play a key role.



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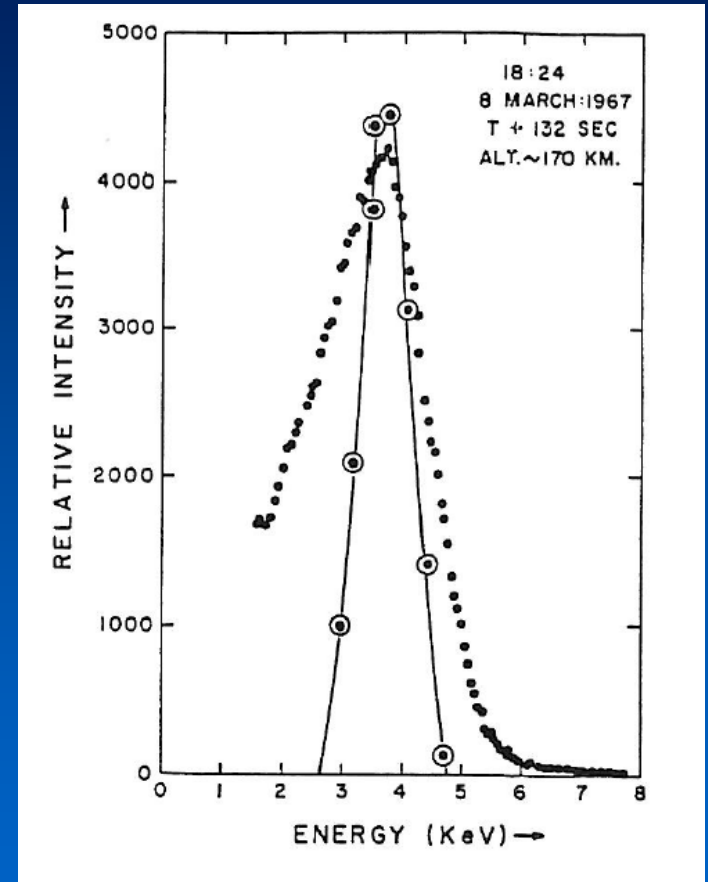
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Early Measurements

- From McIlwain, 1960:
"The presence of monoenergetic electrons strongly suggests an electrostatic acceleration mechanism."



Kletzing et al., [1998]



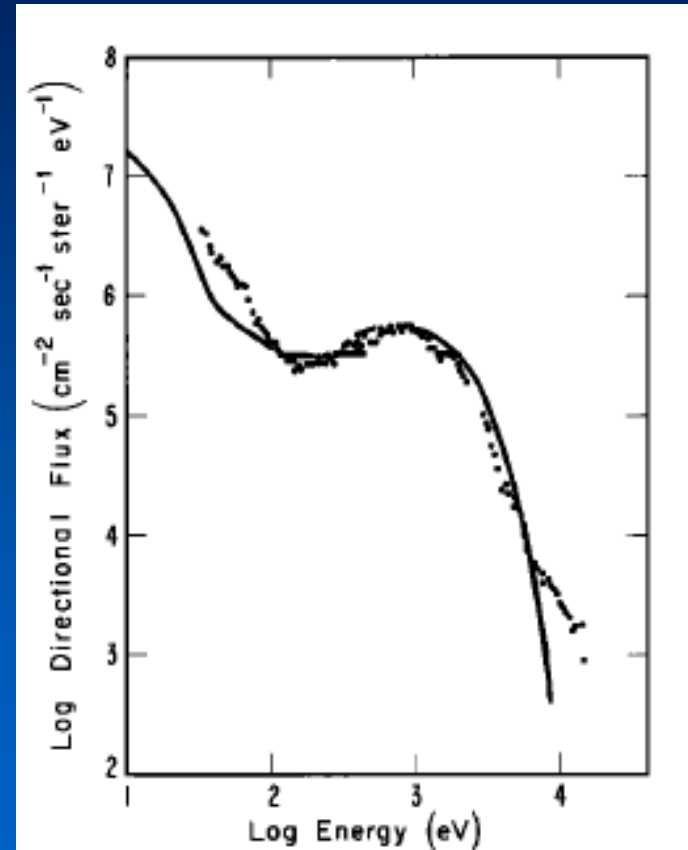


Establishing $E_{||}$

- To get $j_{||}$ to the ionosphere, a potential drop was required.
- Magnetic mirror below and electrostatic mirror above explained observations.

$$j_{||,i}(\Delta\Phi) = -en_m \left(\frac{k_B T}{2\pi m_e} \right)^{\frac{1}{2}} R_B \left[1 - \left(1 - R_B^{-1} \right) \exp \left\{ -\frac{e\Delta\Phi}{k_B T_m (R_B - 1)} \right\} \right],$$

Knight, [1973]



Evans, [1968]





Upward Cuurent Region

1. Upward current region



2. Converging electrostatic shocks.



3. Large-scale density cavity



4. Down-going, "inverted-V" and field-aligned electrons.



5. Up-going ion beams. Ion conics.



6. Large-amplitude ion cyclotron waves and electric field turbulence.



7. Nonlinear time-domain structures associated with ion wave modes.

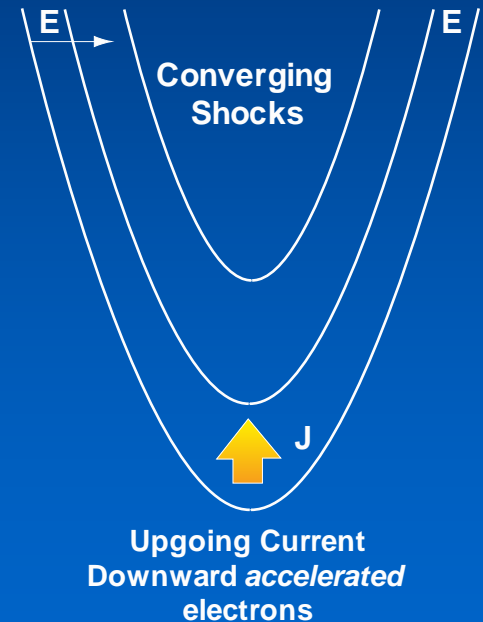
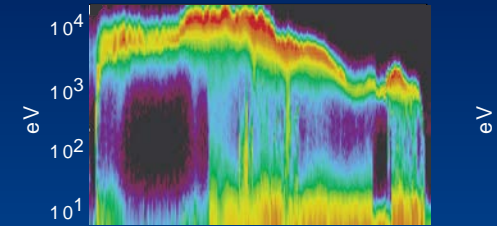


8. AKR source region.



Aurora

Downgoing Electron:





Downward Current Region

1. Downward current region.



2. Diverging electrostatic shocks.



3. Small-scale density cavities.



4. Up-going, field-aligned electrons.
Counter-streaming electrons.



5. Ion heating transverse to **B**.
Energetic ion conics.



6. ELF electric field turbulence. Ion
cyclotron waves.



7. Fast solitary waves: three-dimensional,
rapidly moving electron holes.

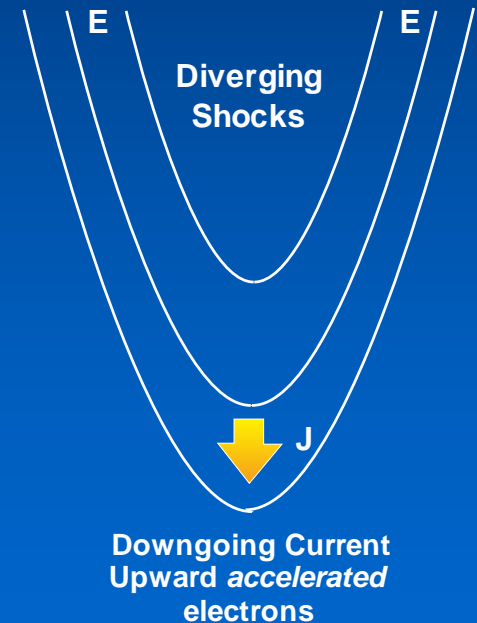
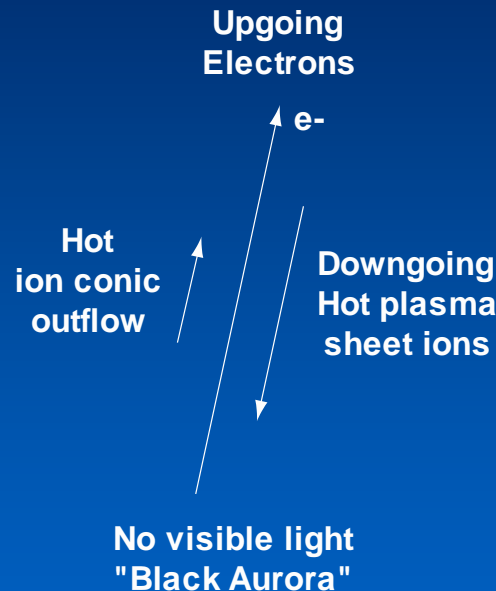
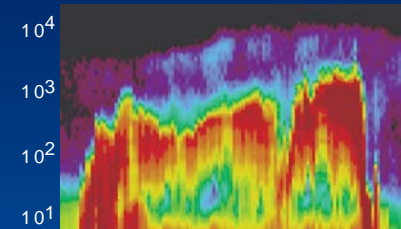


8. VLF saucer source region.



Inverse Aurora

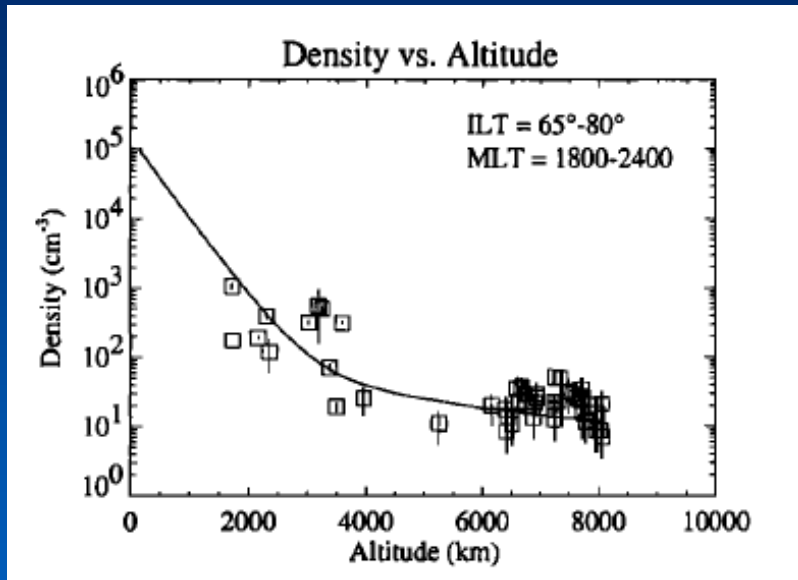
Upgoing Electron:





Densities and Temperatures

Acceleration Region



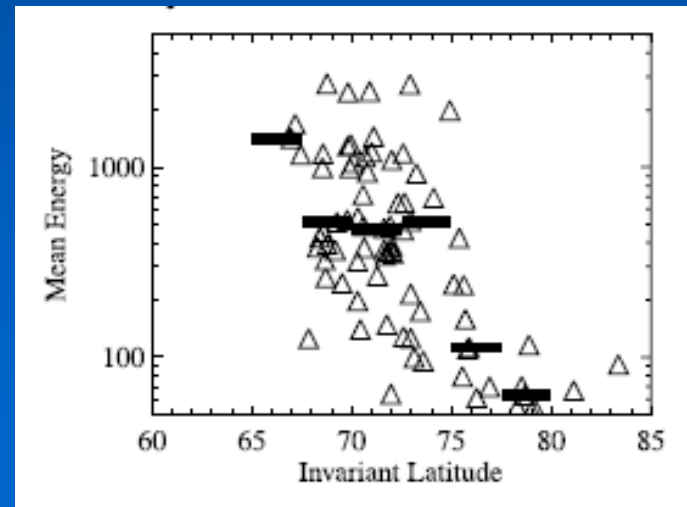
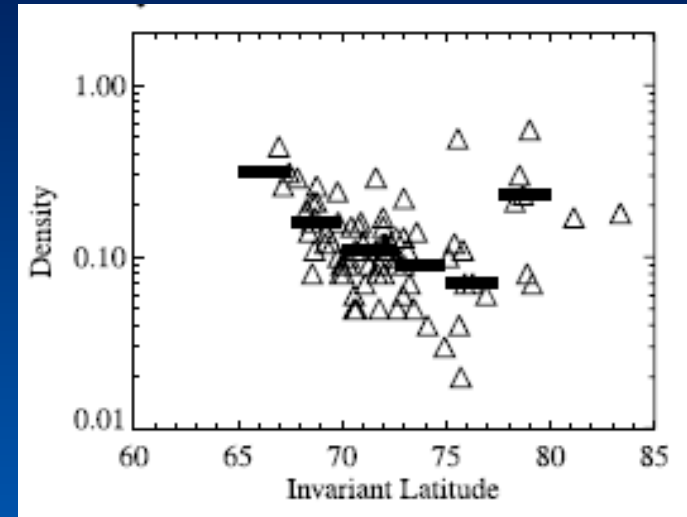
$$n(z) = n_0 e^{-(z-z_0)/h} + n_1 z^{-1.55}$$

Kletzing et al., [1998]



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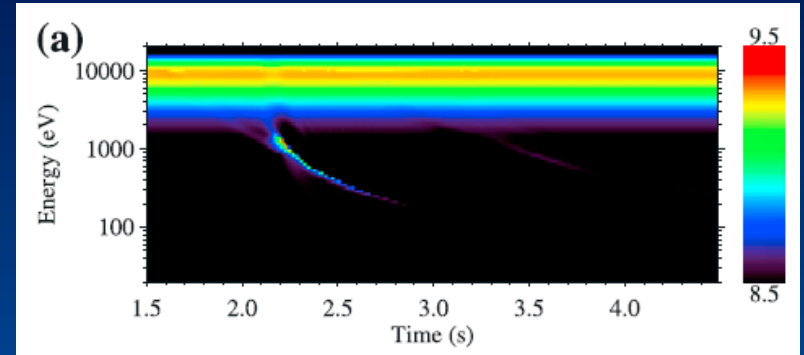
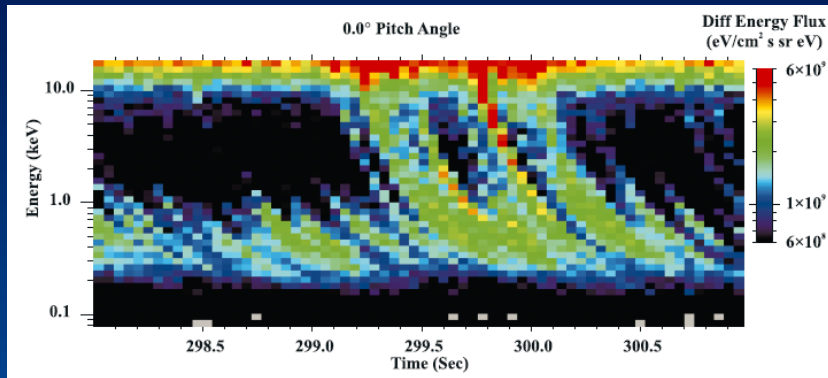
Source Region



Kletzing et al., [2003]

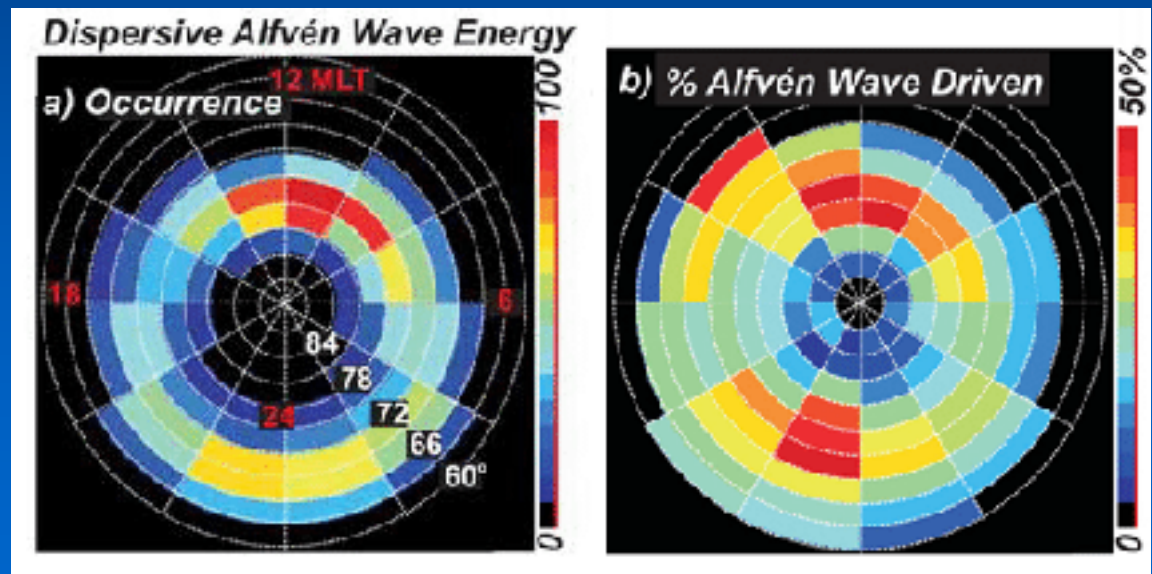


Alfven Waves



Chen et al., [2005]

- Time dispersed electrons indicate Alfven waves.
- FAST survey shows they are ubiquitous.



Chaston et al., [2007]





What We Don't Know

- What is the distribution of the potential drop along the magnetic field?
- How/Do inertial scale Alfvén waves evolve to establish a quasi-static potential drop?
- How do we construct a self-consistent model of the plasma along a field line.
- What are the details of current closure through the ionosphere.



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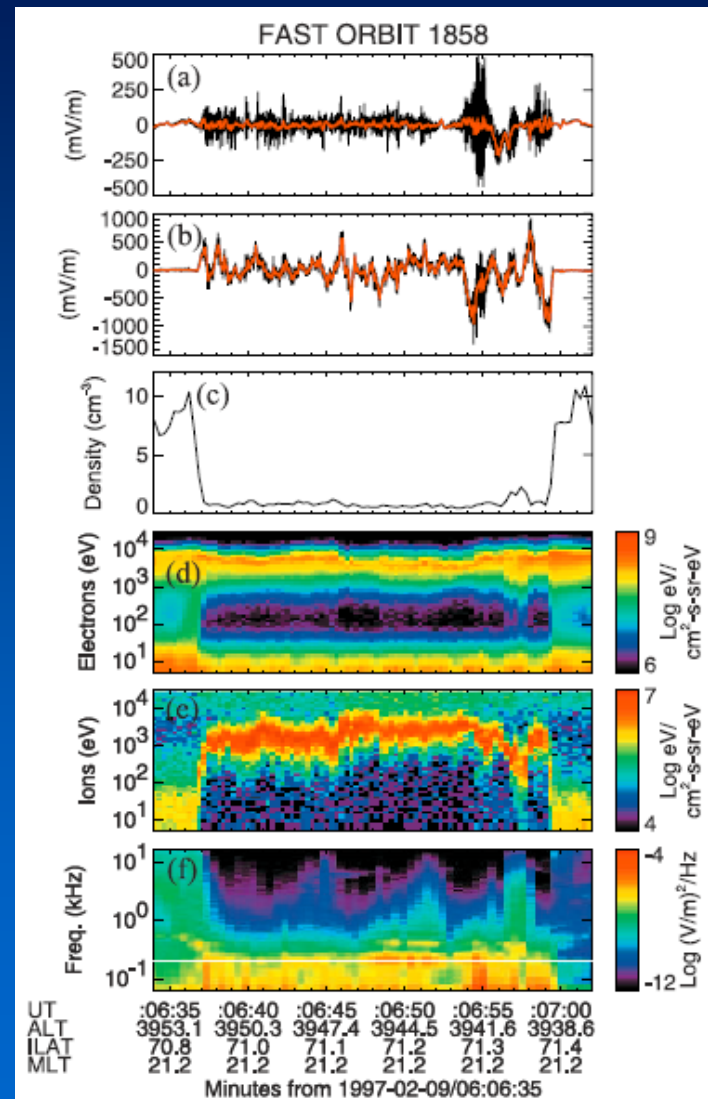
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Potential Along Field Line

- How to determine this?
- Particle distributions: Chiu and Schultz, [1978], Ergun, [2004]
- Direct measurements?



E_{\parallel}

E_{\perp}

$n_i (>5\text{keV})$

electrons

ions

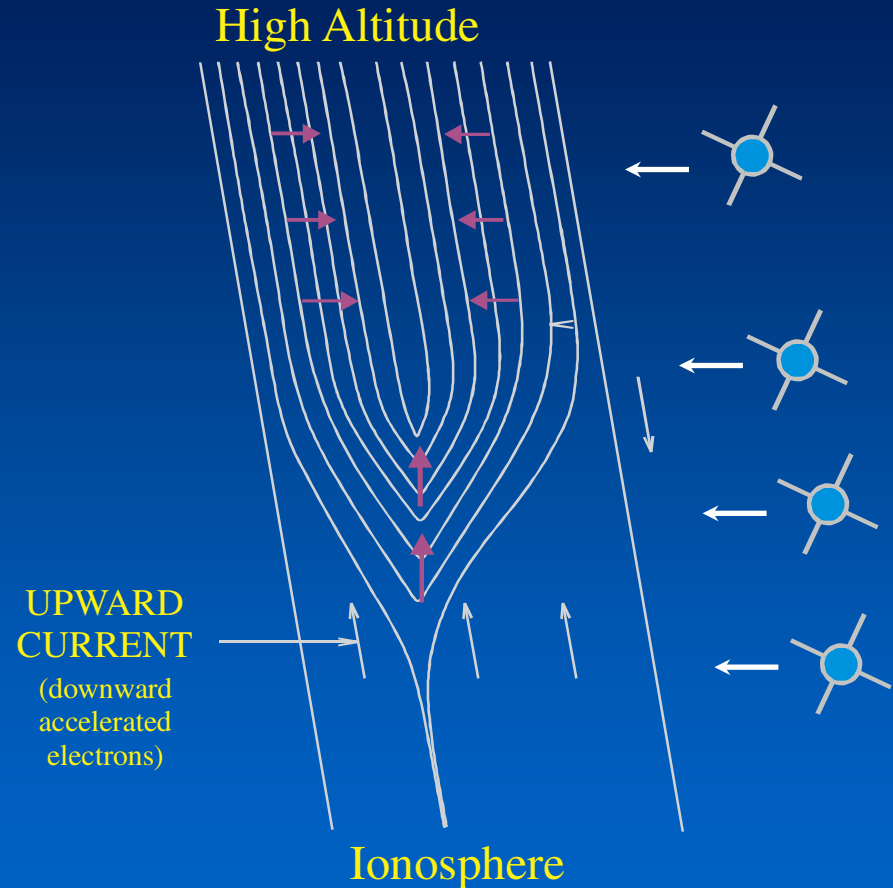
waves





Measuring Potential Drop

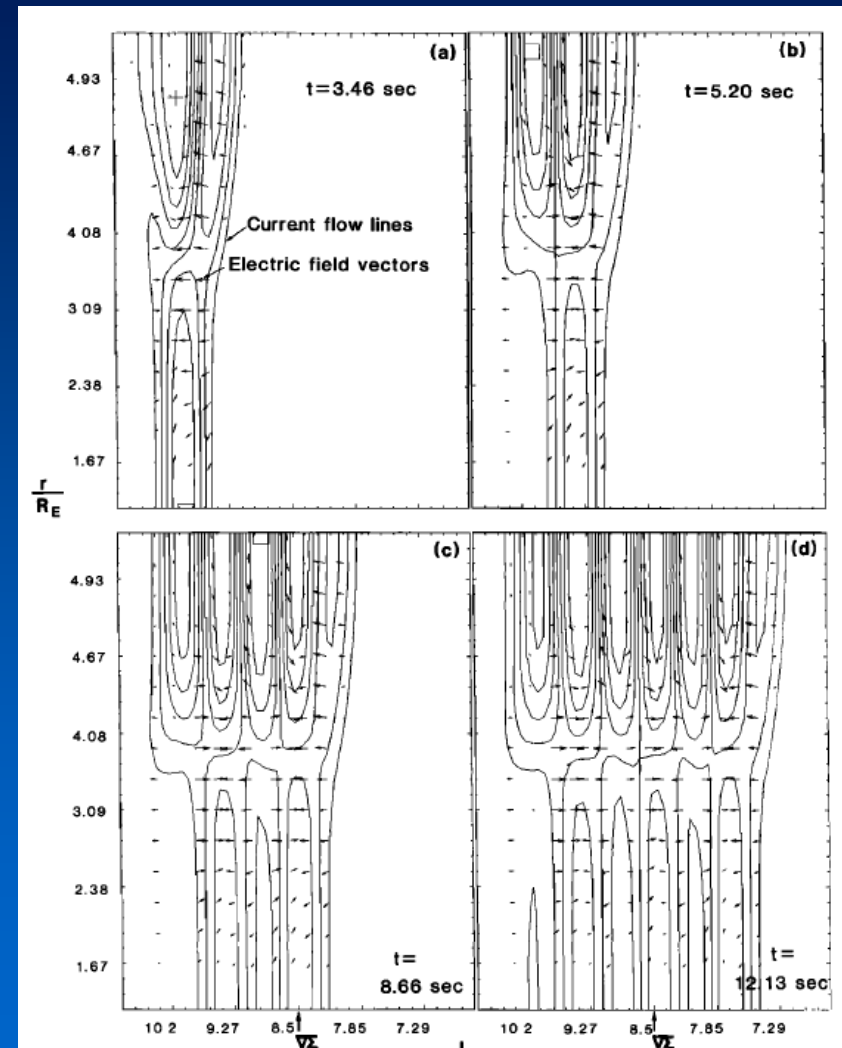
- How to determine this?
- Particle distributions: Chiu and Schultz, [1978], Ergun, [2004]
- Direct measurements?





Alfven Waves & Stable Arcs

- What is the connection?
- Do Alfven waves evolve to 'stable' arcs?
- If so, how does this happen?



Lysak, [1985]

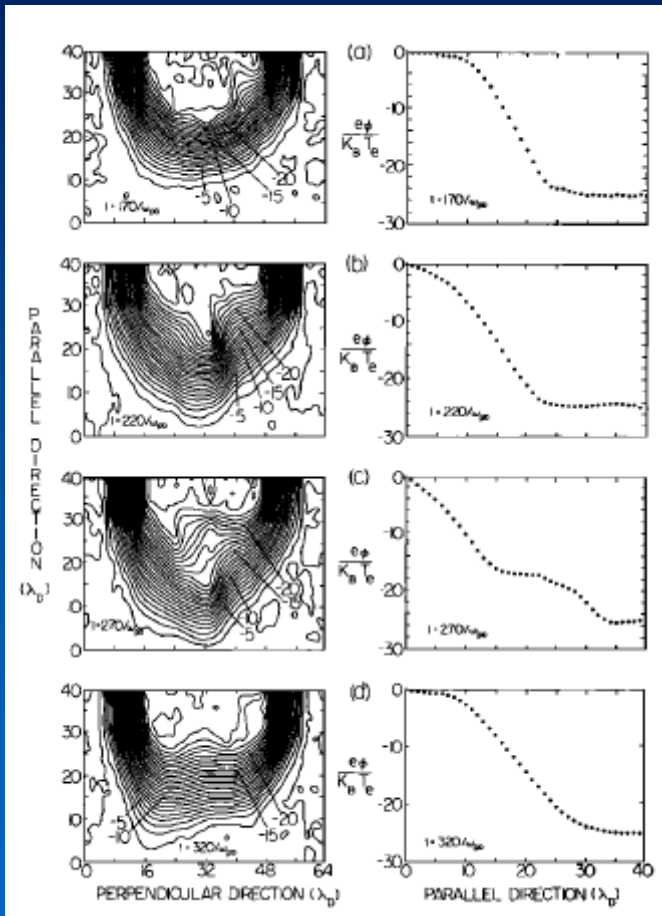




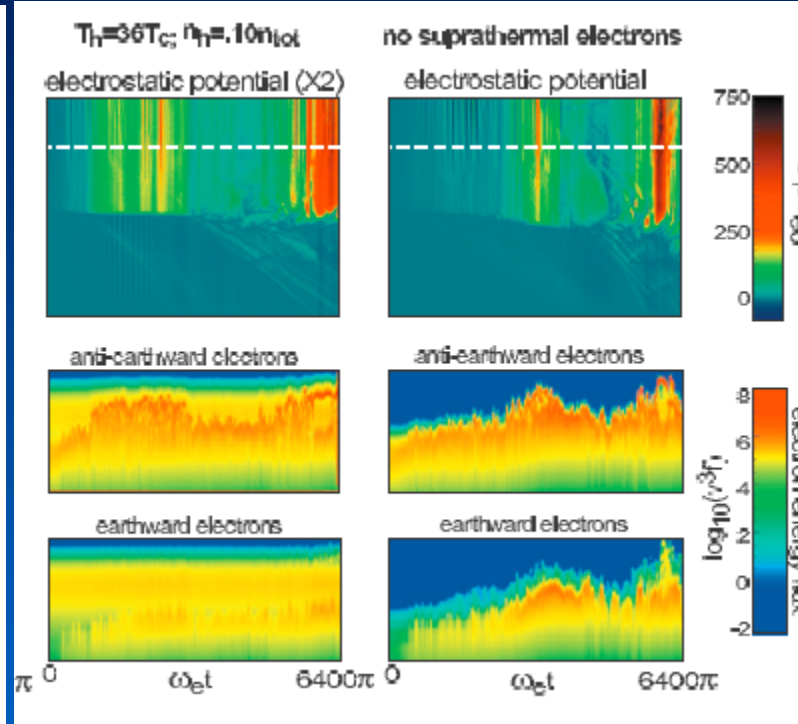
Acceleration Region Model

Then

Now



Borovsky and Joyce, [1983]

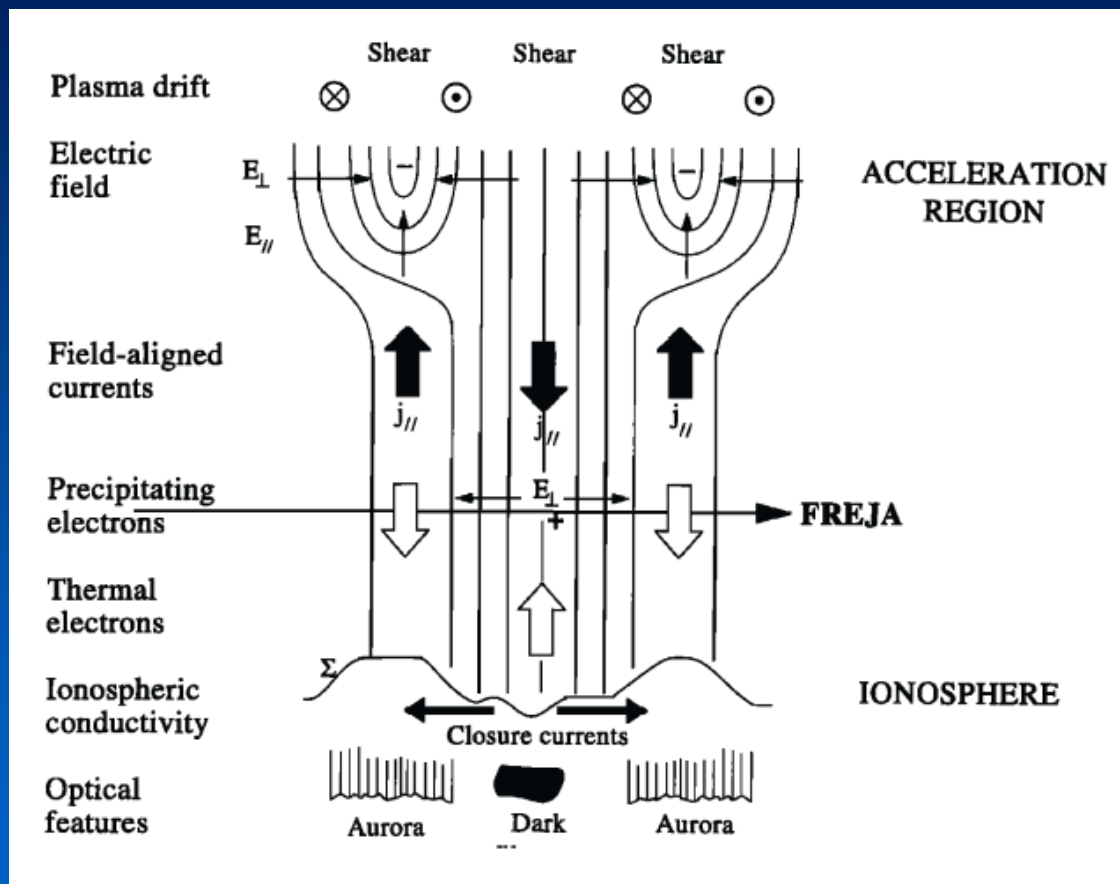


Newman et al., [2008]

- Only small region of field line modelled.
- Not self-consistent.
- Current or voltage driven?



Current Closure



- Can ionization, electric fields and currents be self-consistently modeled?
- Downward region should be much lower density, but rarely observed.

Marklund et al., [1997]



Problems to Work On

- What is the distribution of the potential drop along the magnetic field? – *Need multipoint measurements*
- How/Do inertial scale Alfvén waves evolve to establish a quasi-static potential drop? - *More realistic models needed*
- How do we construct a self-consistent model of the plasma along a field line. – *Large simulations?*
- What are the details of current closure through the ionosphere. – *More measurements and models.*





That's all folks!

