

# What is the origin of low-energy electrons in the inner magnetosphere?

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# A Warm Electron Cloak in the Magnetosphere?

The DENSITY of the overall electron population in the magnetosphere is dominated by the low-energies

*Chappell et al., 2008*

Warm plasma (ion) cloak - 10s eV - effects on dayside reconnection?

Is there an analagous “warm electron cloak” (10s-100s of eV)?

What effects does this population have (if any) on the magnetosphere?

What are the temporal and spatial dynamics of this population?

Wave-effects (e.g. whistler waves?)

Should we care.....?

# Low Energy Electrons

▣ Measuring the low energy (1-100s eV) electron population in the inner magnetosphere is **DIFFICULT** (surface charging - net negative charge in eclipse).

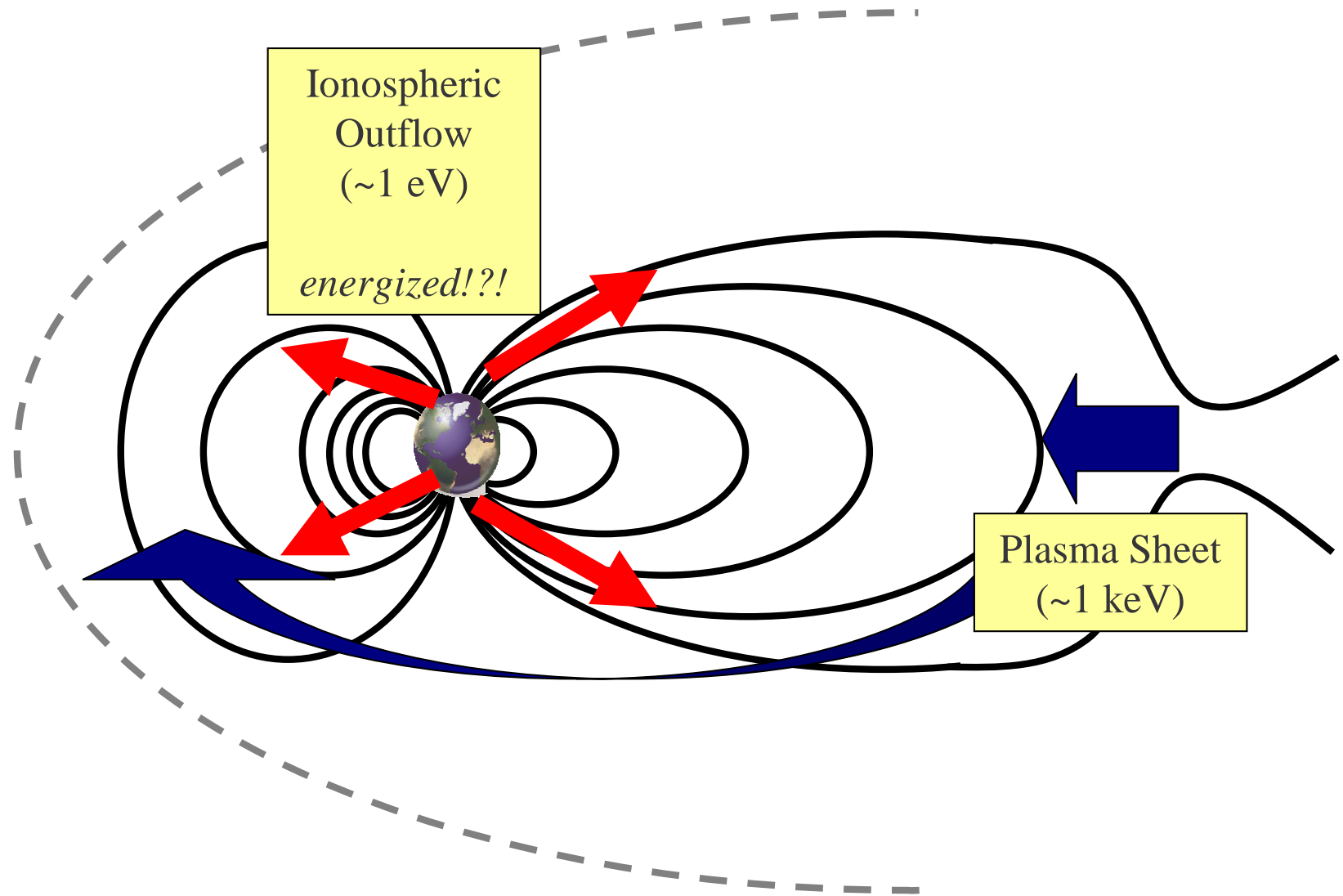
▣ Spacecraft surfaces can charge either positively or negatively - in the absence of sunlight the spacecraft potential will be negative (e.g. *DeForest* [1972], *Garrett* [1981], *Farthing et al.* [1982], *Lanzerotti et al.* [1998], *Thomsen et al.* [2013]).

▣ Electrons with energies below the surface potential are repelled from on-board instrumentation. Any electrons detected will thus be “secondary-electrons” ejected from the spacecraft due to higher-energy particles impinging on the satellite surface.

▣ Electrons with energies from 10s to 100s eV have been poorly sampled in the past.

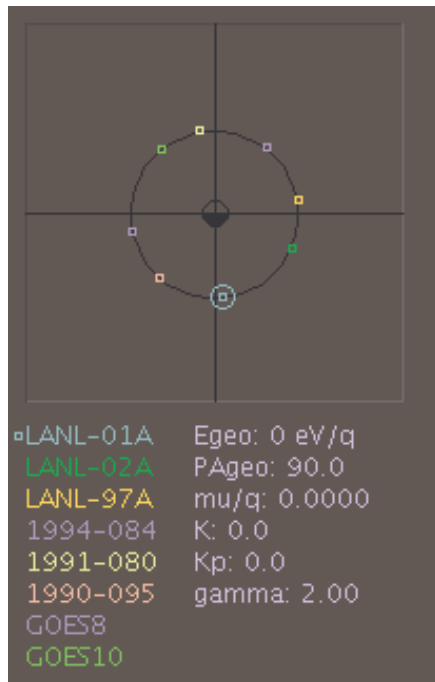
▣ This population, its dynamics and evolution, has been neglected in comparison with other magnetospheric populations such as the plasmasphere, the plasma sheet, the radiation belts, etc.

# Low Energy Electron Sources?



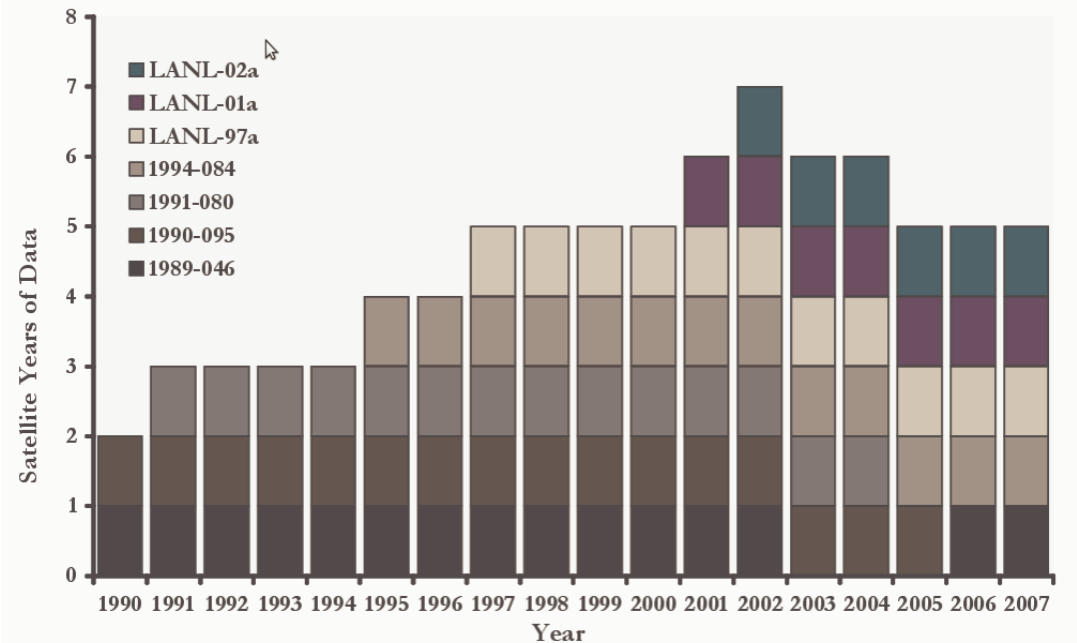
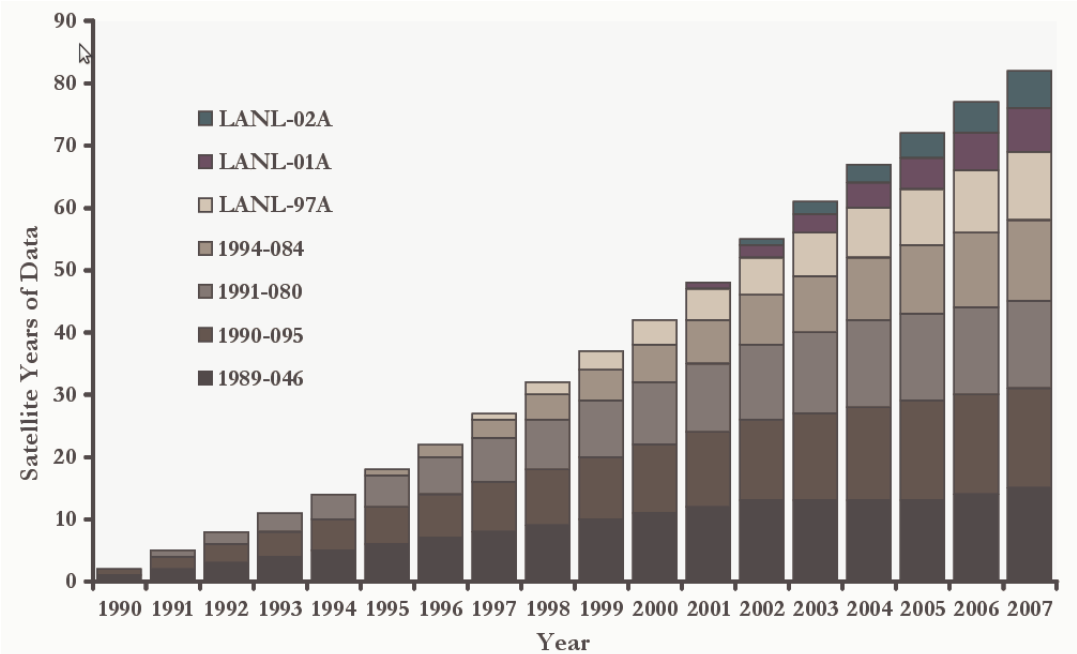
# Observations from GEO

Use LANL/MPA data to explore population statistically

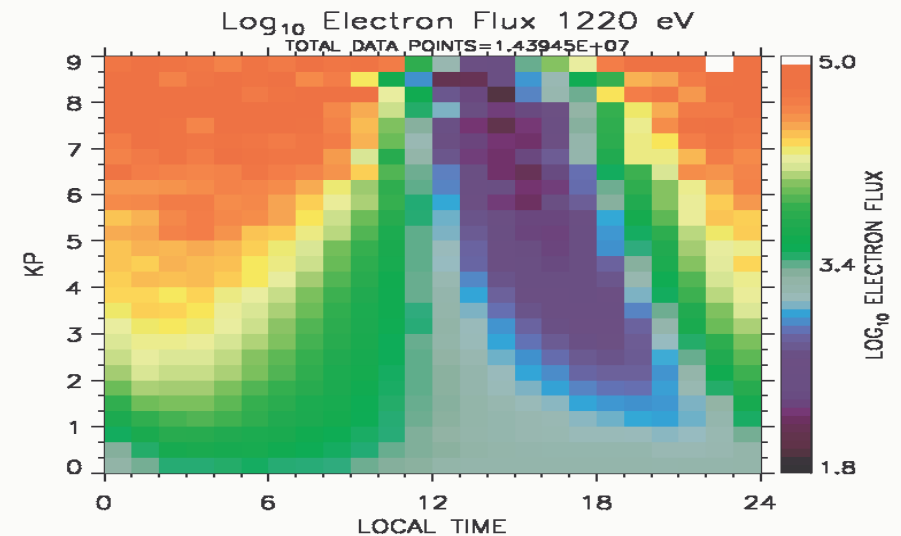
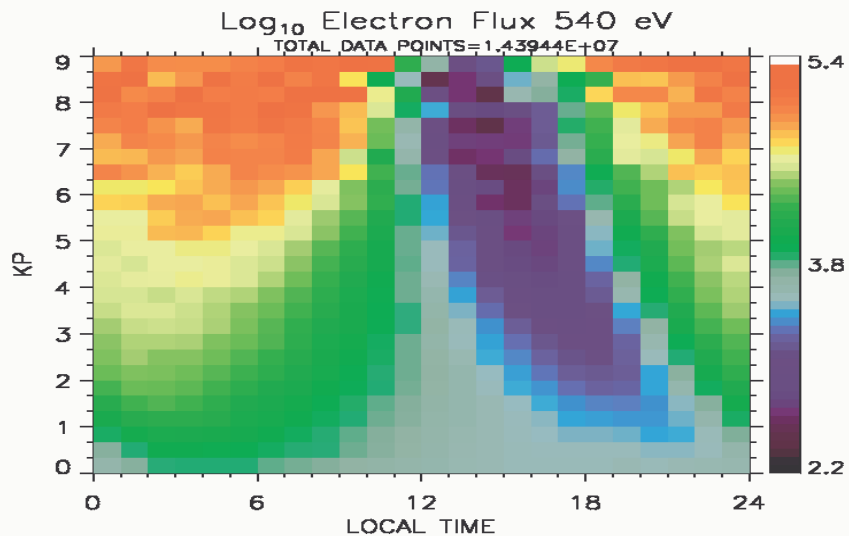
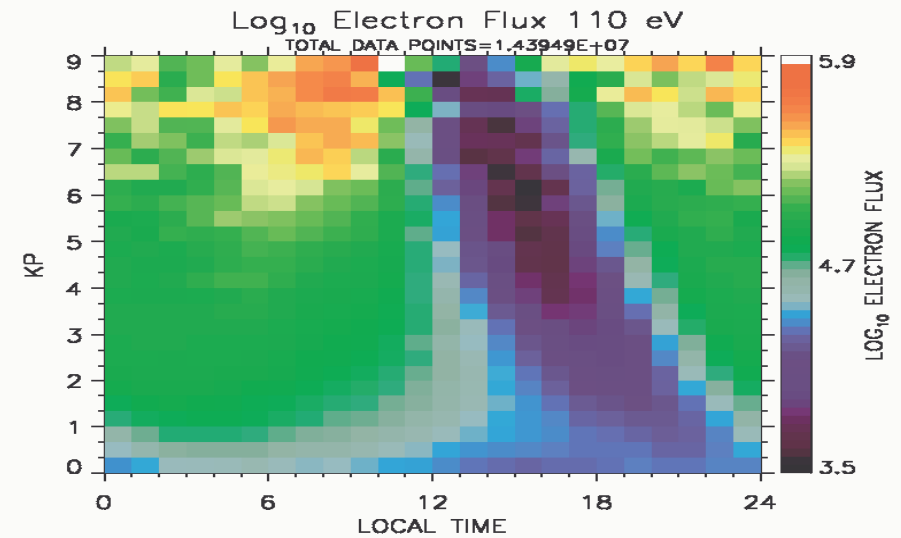
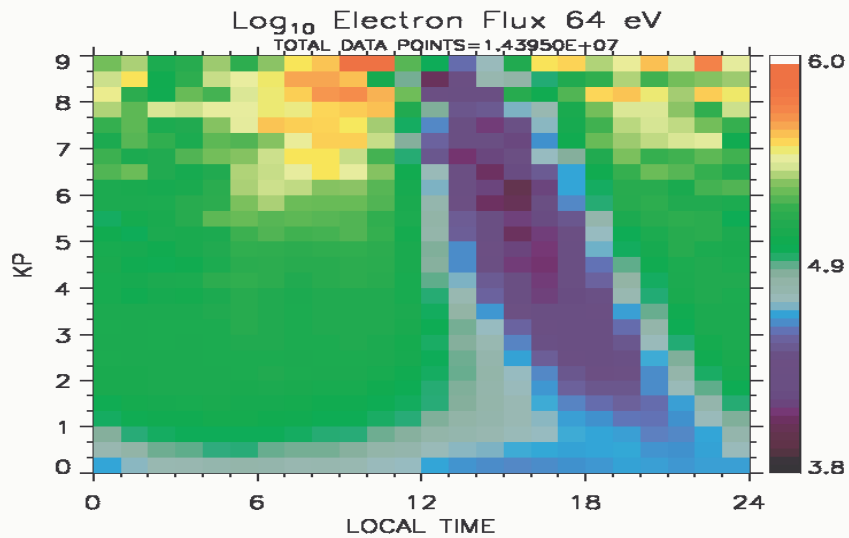


MPA is susceptible to severe negative spacecraft charging

We restrict analysis to times when charging is  $> -40$  V



# Low Energy Electrons



At GEO the electron population with energies 10s-100s of eV is mostly comprised of the low energy tail of the plasma sheet.

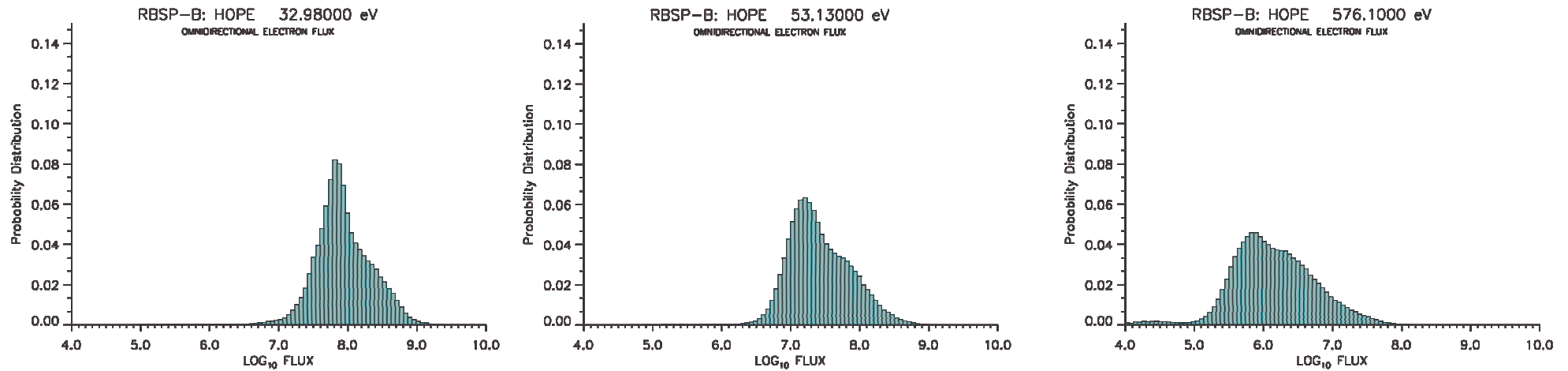
# **RBSP/HOPE**

**For lower energies, we can now use HOPE electron observations to explore inwards of GEO.**

**RBSP designed to minimize spacecraft charging (also it has been a mostly benign environment since launch)**

**From 2013 to 2015 there were only 30 instances where the satellite charged to a potential below -50 V (usually a few volts negative)**

# RBSP/HOPE



Probability distributions of electron flux over full mission (All L, All LT).

Appearance is suggestive of a superposition of two populations...



# RBSP/HOPE

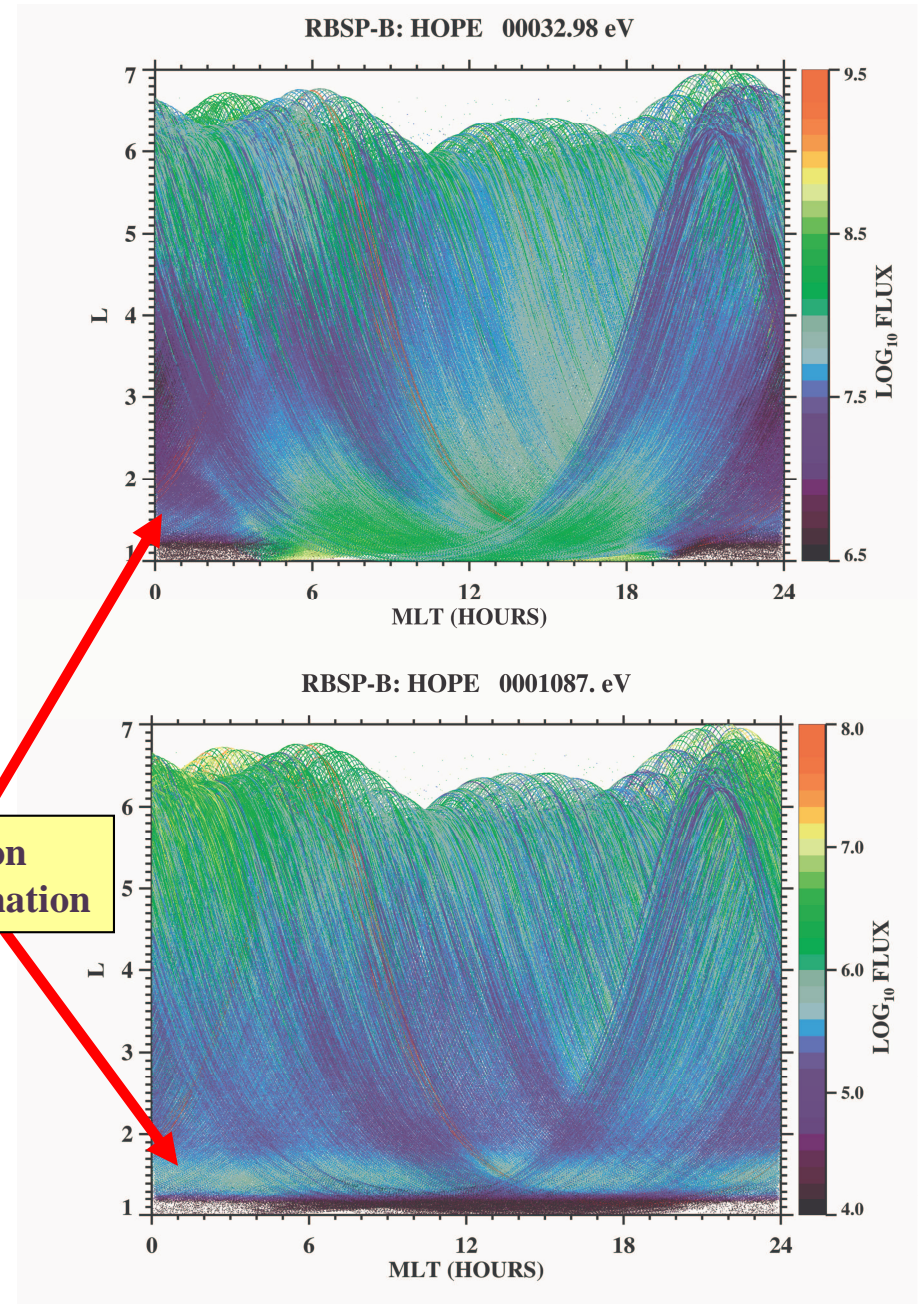
Every (omni-directional) electron data point plotted as a function of L and MLT.

Two populations:

IONOSPHERIC OUTFLOW

PLASMA SHEET ENTRY

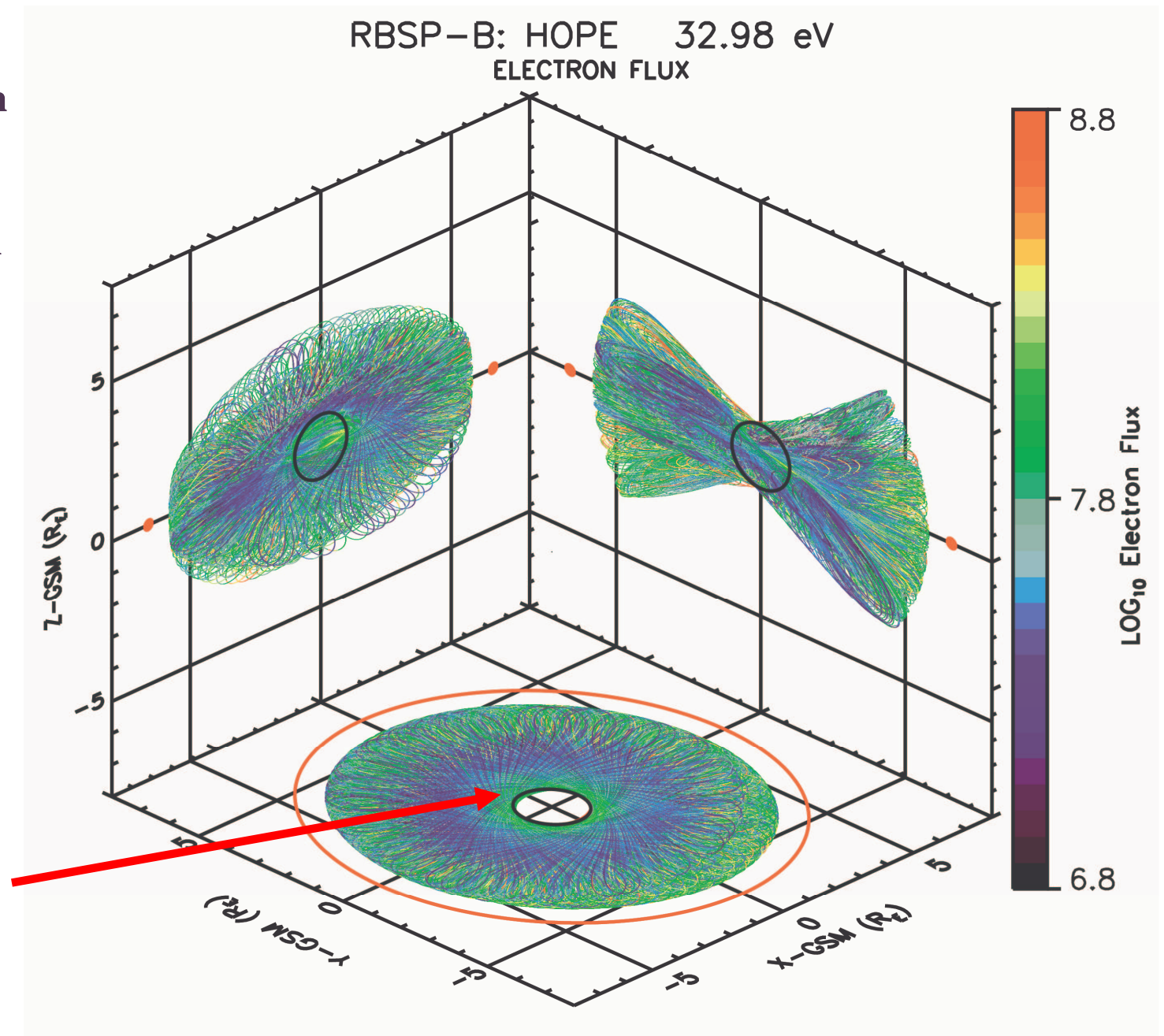
Proton  
contamination



# RBSP/HOPE

Omni-directional  
electrons as a function  
of GSM

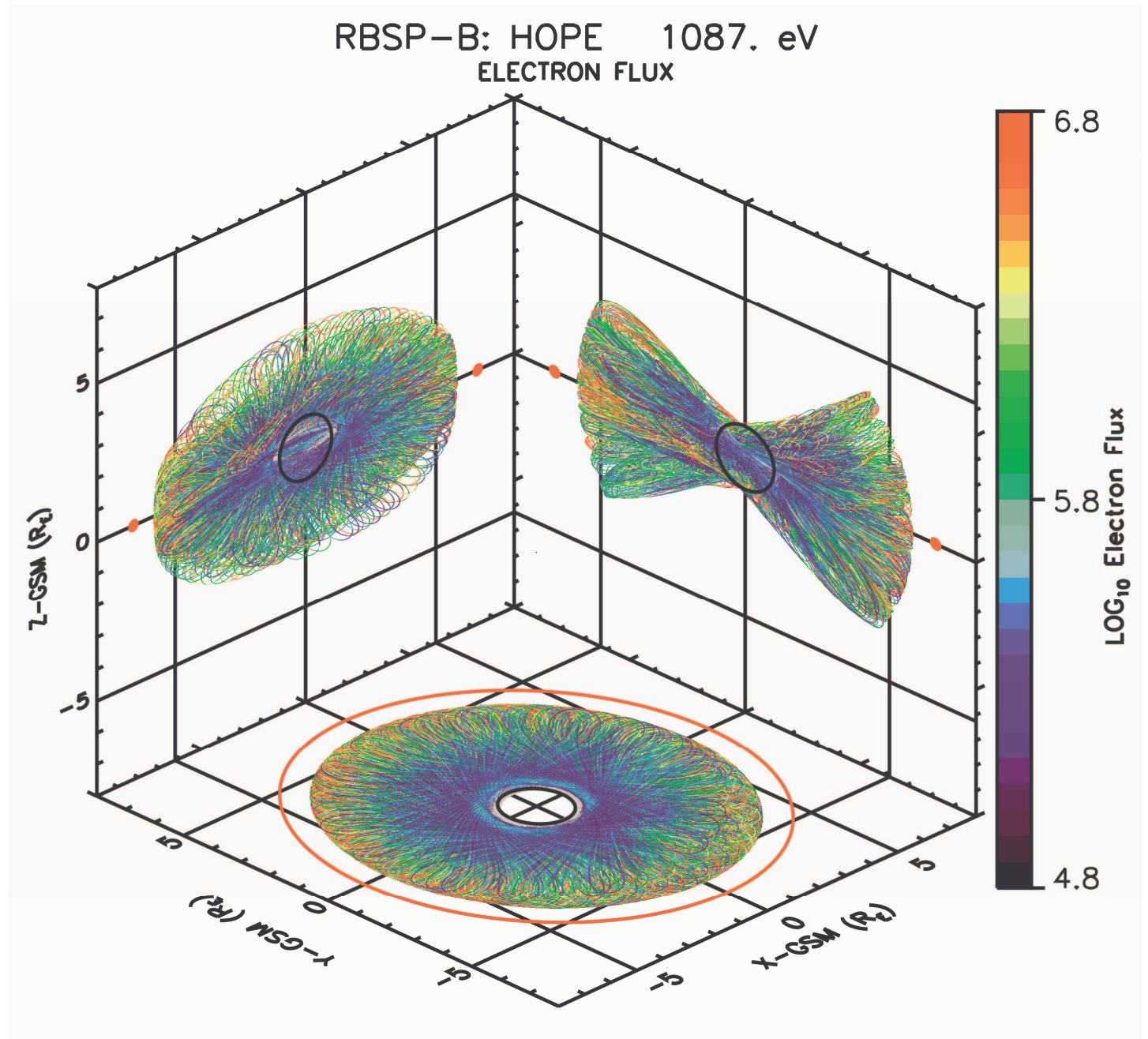
Off-equator variation  
of the population



# RBSP/HOPE

Omni-directional  
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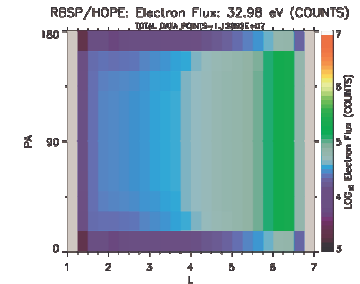
Off-equator variation  
of the population



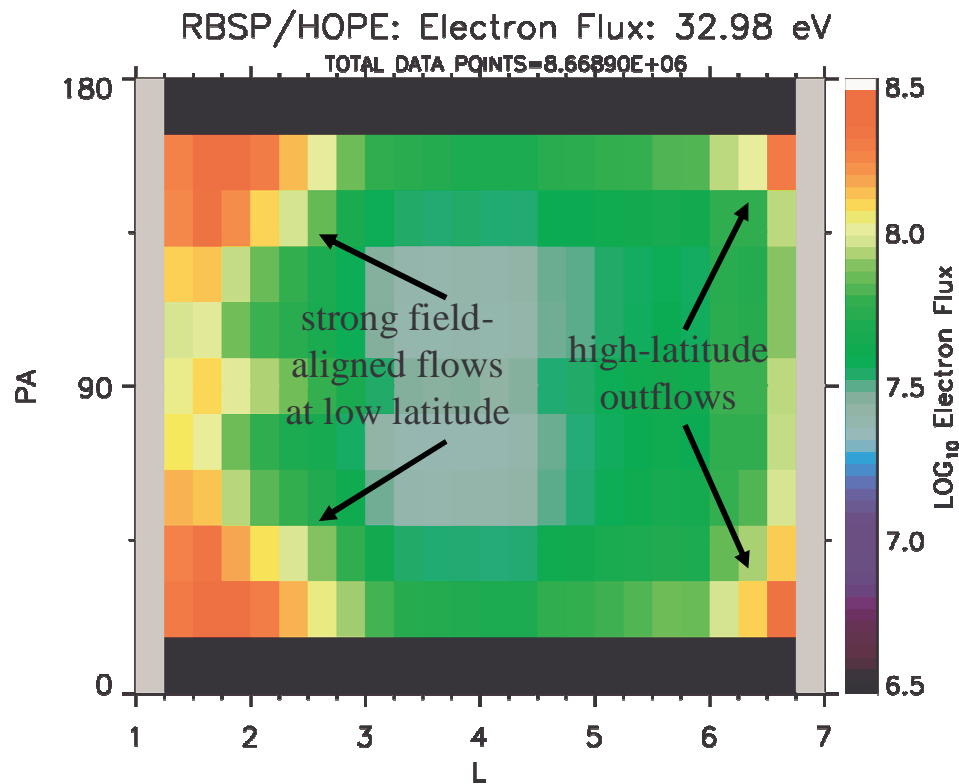
# RBSP/HOPE

Pitch-angle distributions of the electrons sorted by MLT.

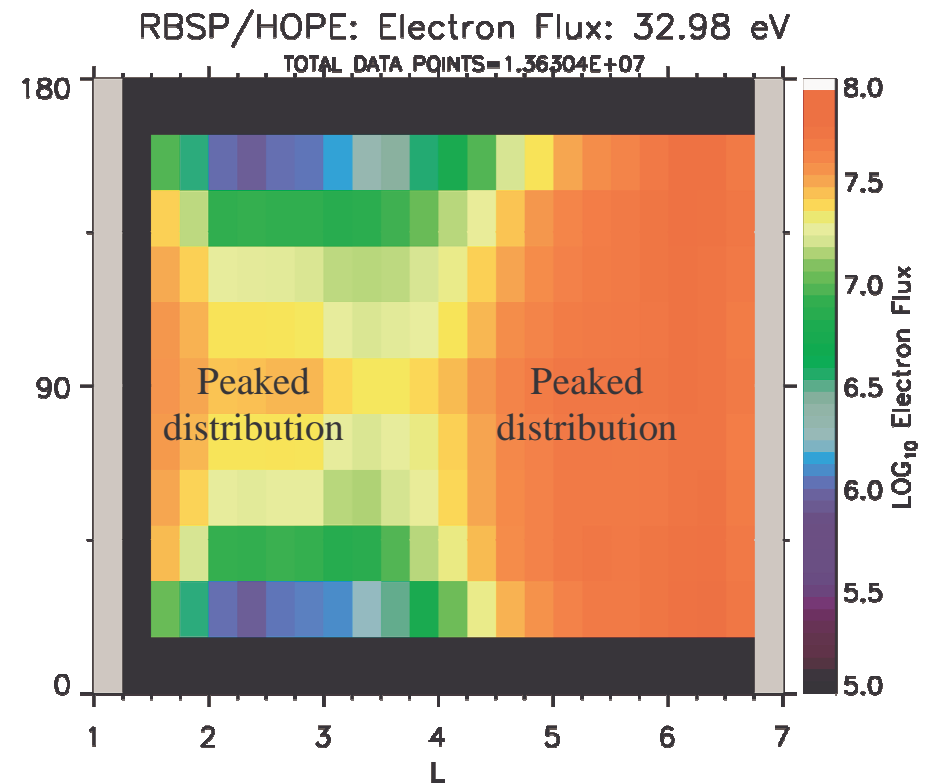
Log (counts)



**NOON (10-14 MLT)**



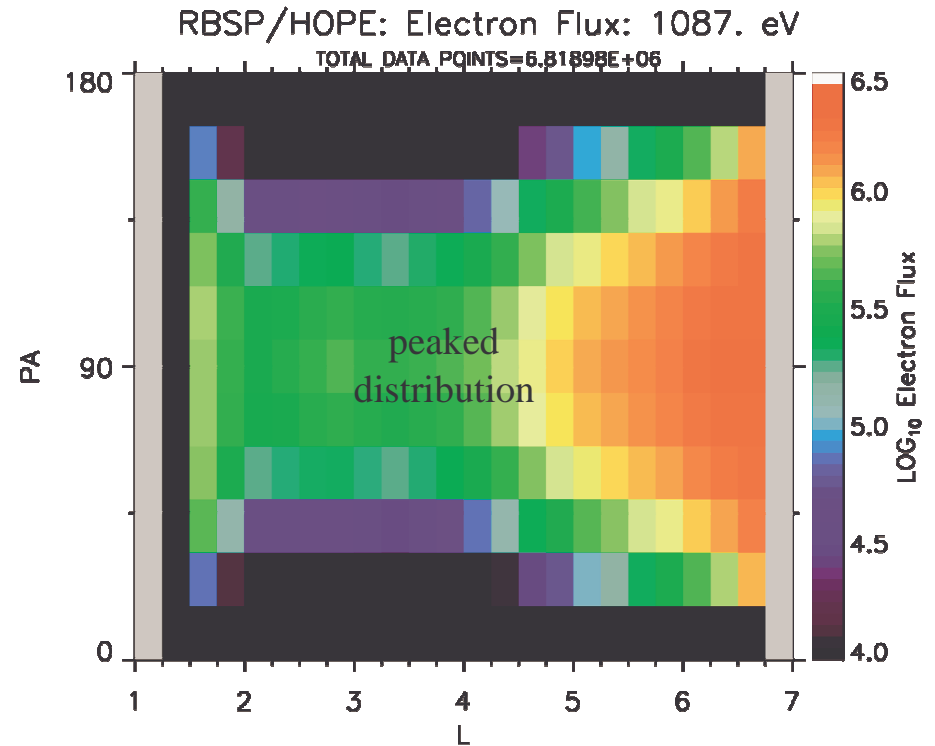
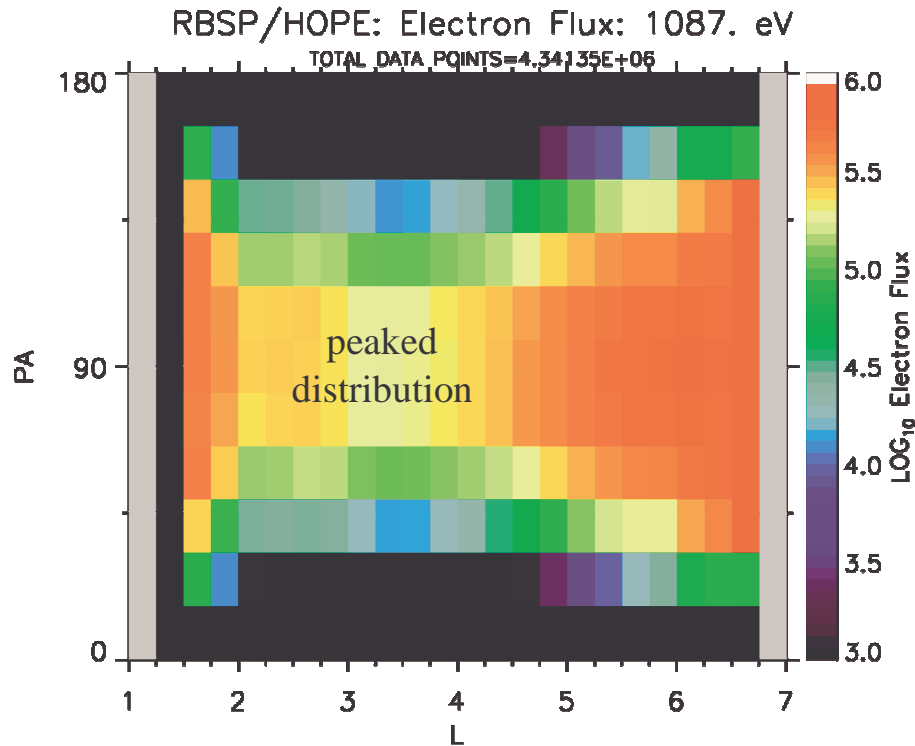
**MIDNIGHT (22-02 MLT)**



# RBSP/HOPE

NOON (10-14 LT)

MIDNIGHT (22-02 LT)



However, this picture is somewhat misleading.

High field-aligned fluxes ARE observed, but not all the time...

# Questions

- ⊙ Can we quantify the conditions that cause field-aligned outflow??
- ⊙ How important is this population for wave growth??
- ⊙ Do we understand the dynamics....should we care??